Suzaku Documentation

Suzaku Users' Group Meeting, 09/21/2009

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on behalf of the

Suzaku GOF @ NASA's GSFC



Documents, Procedures

- Suzaku Data Reduction Guide .pdf, .ps, .html
 ABC Guide .pdf, .ps, .html

http://heasarc.gsfc.nasa.gov/docs/suzaku/
analysis/abc/

Suzaku GOF Web Pages
 http://heasarc.gsfc.nasa.gov/docs/suzaku/astroegof.html

TD, ABC: maintained by Katja Pottschmidt since fall 2008 maintained by Ilana Harrus until fall 2008

WWW: maintained by Koji Mukai

TD: Chapters & General Status

Main document for proposers, **currently being updated for AO-5**, **input provided by instrument teams**, "polished" before AO-4:

- 1 Introduction: *Done*. Minor updates.
- 2 Changes Since AO-4: *Done*. Updated paragraph on the XIS Timing mode (usage possible in AO-5).
- 3 Observation Policies: *Done (Tab. 3.2?)*. Minor updates.
- Writing A Successful Suzaku Proposal: To Do, Katja: Simulations with new HXD responses and backgrounds. Draft files provided (HXD team).
- 5 XRT: *Done*. No updates required (XRT team).
- 6 XIS: *Done*. Revised chapter & plots provided (XIS team).
- **THATE:** Done. No updates, GSO in prep. (HXD team).

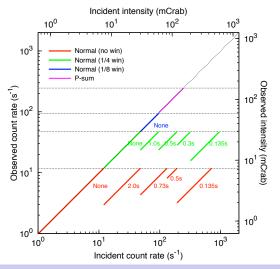


TD: XIS Changes

- 7.5.3 Editing Modes:
 "Note on the Timing mode" rewritten and expanded.
- 7.6.2 Supported Modes (for SCI):
 New tab. "Supported clock modes and options with SCI".
- 7.9.1 Photon Pile-Up:
 New figure "Incident vs observed count rates of a point source for the FI sensor". New statement: "ASCA grade 1 fraction correlation with pile-up fraction".
- 7.10 Clock/editing Mode Selection:
 New table "Maximum incident counting rate for avoiding pile-up, observational efficiency, and maximum observed counting rate". Replaces two old tables.
- 7.12 Anomaly in XISO and XIS2
- 7.13 XIS Quick Reference: New figure.



TD: XIS Changes







Suzaku X-ray Imaging Spectrometer Quick Reference

2009/08/13 M. Tsujimoto (ISAS)

This leaflet is intended to assist novice users to plan a XIS observation. The Suzaku web page (http://www.astro.isas.jaxa.jp/suzaku/index.html.en) and the "Technical Description" document supplement the information. Consult xishelp@astro.isas.jaxa.jp for further details.

accuracy, ...

Basics XIS is equipped with four X-ray CCDs (XISO-3) for imaging and non-dispersive spectroscopy. The four CCDs are at the focus of four co-aligned telescopes and observe the same field. Three CCDs are front-illuminated (FI) and one is back-illuminated (FI) show the first four first four first four first four first four first first four first fir

Counts/s

Estimate the count rate using the PIMMS tool. Approximately,
Inche hix yeldes 1,6 (ent/s) (F1) and 15 (ent/s) (R1). For point
sources with >12 (ent/s), choose an appropriate clocking mode to mitigate pile-up.

Modes

Users are responsible to choose the clocking mode. The XIS
operation team choose the celliting mode. It is acceptable to use

XIS is operated simultaneously with HXD. Time resolution \$8 \text{ (Normal)} - \text{ } 8 \text{

View XISO-3 has 1024x1024 pixels composed of four segments (A-D) with one readout node for each segment. A part of XISO and the entire XIS2 are not usable. Two ³⁶Fe calibration sources (Mn I Kα and Kβ lines at 5.9 and 6.5 keV) are installed. Users can soedfy the roll anale. Use the Maki tool.

 Cock mode
 *** Support

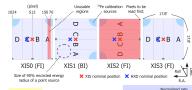
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different clocking modes for different CCDs. Select the normal mode with no

for bright sources, and the Psum mode for extremely bright sources.

options for faint sources, the normal mode with the burst and/or window options



*1: Calibration not guaranteed. Prior consultation required. *2: Only for NiOn-normal position. *3: Only for FI.

<u>Window option</u>... 1/n (n=4 or 8) option reads (1024x1024/n) pixels around the
aim position in 8/n [s]. The size in the readout direction is reduced. (Pros) Photons
not lost for the observed area. (Cons) The observed area reduced by 1/n. The
calibration sources not observe.

Busts aption ... m [s (m=0.1, 0.3, 0.5, 0.7, 2.0) option reads photons arriving in m out of 8 [s] in each image. (Prox) The calibration sources observed. The observation area not reduced. (Cons) A fraction (1-m/8) of photons lost. Bust m code ... 12 frow stadded along the readout direction, yielding (102468) pixel data. (Prox) High timing accuracy of 2.8 ms. (Cons) Spatial information lost reduction. The unabsolute of m considerable m considerable

Others

Simulate your observation. Things to consider include: (1) the increasing contamination materials for soft sources, (2) the non X-ray background for low-surface brightness sources, (3) telemetry asturation for brightness records.

Aim point Choose either XIS- or HXD-nominal position, depending on which detector you emphasize. The count rate differs by ~10%. Other positions cannot be chosen.

 Position
 XIS
 HXD

 XIS nominal
 1
 0.9

 HXD nominal
 0.9
 1



ABC: Chapters & General Status

Current: Version 2, January 2008

In prep.: Version 3

Updated:

- Introduction
- Software
- Suzaku Data Specifics and Conventions
- 4 Suzaku Data Analysis Overview (add flowcharts?)
- 5 The "README FIRST" of Suzaku data analysis
- Strategies
 XIS Data Analysis update on-going
- THXD Data Analysis update on-going, incorporated e-mail from Terada-san



ABC: Changes Since Version 1

- Version 1 ⇒ 2 for the ABC Guide =
 Version 1 ⇒ 2 for the Processing Pipeline
- Version 1 mainly described the processing up to screening event files (i.e., "Initial Processing", "Standard Screening")
- No "Analysis Overview" chapter, minimal "README FIRST" chapter.
- 54 ⇒ 77 pages

Change Log Processing (V1 \Rightarrow V2)

SCI-on to be processed, thermal wobbling correction, P-sum mode to be processed, HXD events FITS format changed, lower PI *E*-threshold revised, GSO *E*-scale calibration revised, data screening criteria updated, etc.

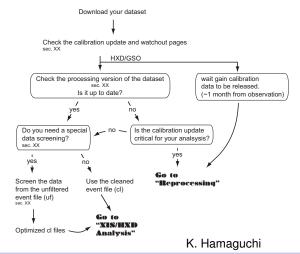
ABC: Updated Chapters

- Chapter 3: Data Specifics and Conventions: updated XSELECT default parameters
- Chapter 4: Data Analysis Overview: added GSO, discussing "flow charts"
- Chapter 5: README FIRST:
 Major update, incorporating many topics from the GOF web pages:

updated: XIS contamination, updated energy scale of SCI-on data, PIN bkg, GSO bkg, cross calibration, etc. new: XISO anomaly, energy scale for window option data, effective area of the XRT/XIS system, XIS pile-up, XIS vignetting correction, HXD reprocessing, etc.

ABC: Recent Work, Draft Flowchart

Need to Reprocess?





Suzaku Data Analysis Web Pages I

For most up-to-date information, check:

- Data Analysis: http://heasarc.gsfc.nasa.gov/ docs/suzaku/aehp_data_analysis.html
 - Things to Watch Out For http://heasarc.gsfc.nasa.gov/docs/suzaku/ analysis/watchout.html
 - Suzaku FTOOLS Version 12
 http://heasarc.gsfc.nasa.gov/docs/suzaku/
 analysis/suzaku_ftools.html
 - Monthly Update
 http://heasarc.gsfc.nasa.gov/docs/heasarc/
 caldb/suzaku/
- What's New: http://heasarc.gsfc.nasa.gov/ docs/suzaku/aehp_whatsnew.html



Suzaku Data Analysis Web Pages II

http://heasarc.gsfc.nasa.gov/docs/suzaku/

- GSO background page and reprocessing instructions: analysis/gsobgd.html, analysis/hxd_repro.html
- PIN background pages and Epoch 5 information: analysis/pinbgd.html, analysis/pin_cxb.html
- Vignetting-corrected XIS images: analysis/expomap.html
- XRT+XIS effective area: analysis/xrt_update.html
- E-scale XIS window: analysis/xis_window.html
- Contamination evolution: analysis/xis0.html
- ISIS tools for pile-up correction: http://space.mit.edu/ASC/software/suzaku/



Things to Watch Out For Last Updated Jun 4, 2009

This page contains the brief summaries of important issues that all Suzaku users should be aware of. It will be updated frequently to reflect our latest understanding of real and potential issues in Suzaku data analysis.

Most V1.x to V2.x transition issues are believed to have been solved.

M (http://heasarc.gsfc.nasa.gov/docs/suzaku/analysis/watchout.html

- All Suzaku observations have been (re)processed with V2.x.
- HXD/PIN background files have been released (updated 2008 Jun 17).
- HXD/GSO background files have been released.

 YIS sight party background database and tool be
- XIS night earth background database and tool have been released.
- There is a new version of the ABC guide that corresponds to V2.x data anilysis, although it does not yet include the description of GSO data analysis.

The remainder of this page documents the known calibration and software issues.



suzaku gof



The remainder of this page documents the known calibration and software issues.

1. Calibration Issues

ISAS maintains the calibration status and caveats pages.

In addition, we note the following items:

- 1. HXD/PIN XIS Cross Normalization (Updated 2008 Jul 31)
 - With the 2008-07-09 release of CALDB used with Version 2 data, the normalization of PIN data relative to XISO data is 1.16 for observations of the Crab at the XIS nominal position, and 1.18 for those at the HXD nominal position (uncertainty is -0.015). See Suzaku Memo 2008-06 for details.

With older calibrations, the normalization of PIN data relative to XISO data is 1.06-1.09 for observations of the Crab at the XIS nominal position, and 1.11-1.13 for those at the HXD nominal position. The rangees are for use of different energy bands for the PIN data. See Suzuaku Memo 2007-11 for details.

- PIN GSO Cross Normalization (Posted on 2008 Jun 27)
 Mills the approach as liberation in the 2007 March 20
 - With the current calibration, in the 2007 March 20 observation of the Crab at the HXD nominal position, there is a cross-normalization problem at the 20% level (1.0:0.80, PIN:GSO) and ~10% residuals in the GSO data.

The HXD team is working on resolving this issue. In the mean time, $\underline{\underline{\underline{\underline{smpirical arfs}}}}$ can be used to reduce both problems significantly.

- 3. XIS: Energy scale and resolution calibration as verified using the $^{55}{\rm Fe}$ calibration soruce
 - O SCI data processed using V2.16.15 and earlier versions: After 2006 September, the gain at Mn K alpha line decreased at a rate of 30 eV/year in the FI chips and 50 eV/year in the BI chip. See notes on V2.0/2.1 processed XIS data for more. Also read about updated calibration.
 - Non-SCI data: The energy scale is well calibrated untill July 2006 in FI units, but systematically higher by 10 eV in XIS1.
 After this time, gain gradually increased in XIS3, rapidly decreased in XIS0 and XIS2, and gradually decreased in XIS1.
 - 2 x2 mode (non-SCI) (Updated 2008 Oct 22): There was an offset in the 2x2 mode energy scale relative to the 3x3/5x5 modes. This is believed fixed with the release of ae_xiN_makepi_20080825.fits files in 2008 September. Users should apply the new calibration using the same procedure as for SCI-on data processed with V2.16.15 or earlier.
 - 2x2 mode (SCI-on): (Added on 2008 Oct 22): The energy scale for 2x2 mode with charge injection is expected to be similar
 to those of the 3x3/5x5 modes. However, this expectation has not yet been verified with actual calibration. Users of 2x2
 mode data with SCI should carefully check for gain discrepancies.
 - O Window mode data: (Added on 2009 Jun 4): The energy scale calibration for XIS data with the 1/4 and 1/8 window mode

