A Suzaku Observation of MCG-2-58-22

Constraining the Geometry of the Circumnuclear Material

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Outline

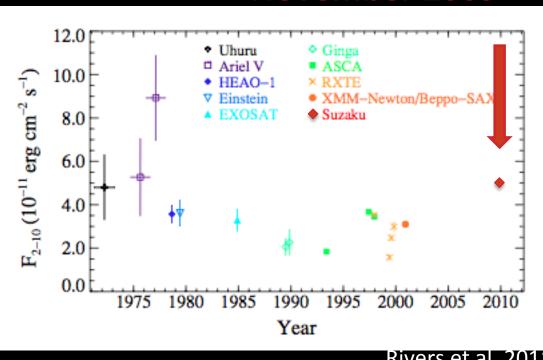
- MCG-2-58-22 Background
- The Fe K Complex
- Reflection Models
 - Compton Reflection from the Accretion Disk
 - The MYTorus Model
- Summary

MCG-2-58-22

- AGN Type: Seyfert 1.5
- X-Ray Components:
 - Unabsorbed
 - Fe K Lines
 - Soft Excess
 - Compton

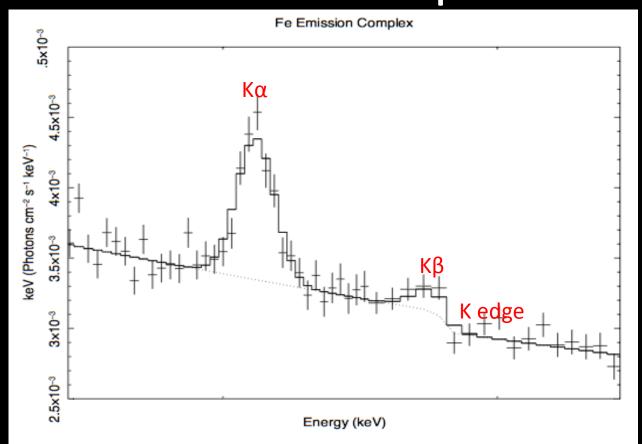
Reflection Hump

140 ks observation **November 2009**



Rivers et al. 2011

The Fe K Complex



- No broad line, Narrow Kα line
- Strong K edge, Fe Kβ/Fe XXVI line

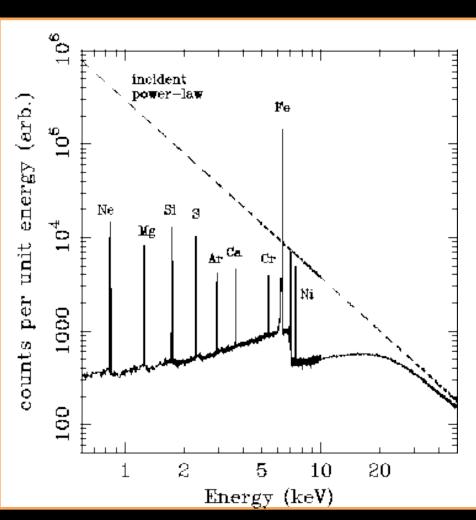
The Fe K Complex

• Weak narrow $K\alpha$ line: $EW = 50 \pm 10 \text{ eV}$

• Width: $v_{FWHM} \le 7100 \text{ km s}^{-1}$ $M_{BH} = 10^{8.4} \text{ M}_{\odot}$ $\rightarrow R \ge 45 \text{ lt days } (\approx 1200 \text{ R}_S)$

• BLR: H β v_{FWHM} \approx 6400 – 8600 km s⁻¹

Compton Reflection off a Disk



Compton scattering off of Compton-thick material $(N_H^{\sim} 10^{24} \text{ cm}^{-2})$ Produces emission lines and Compton Reflection Hump (CRH)

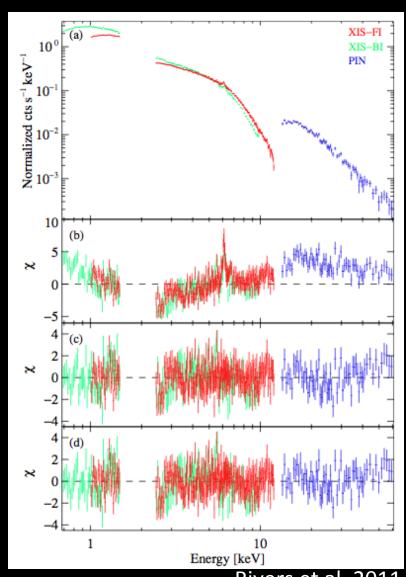
CRH Strength - R

Reynolds & Nowak 2003

Disk Modelling Results

- $\Gamma = 1.80 \pm 0.02$
- $\Gamma_{SX} = 3.0 \pm 0.6$
- $R = 0.69 \pm 0.05$

Fe line EW slightly lower than expected from CRH



Rivers et al. 2011

Compton Reflection and Reprocessing by a Dusty Torus

The MYTorus Model

(Murphy & Yaqoob 2009)

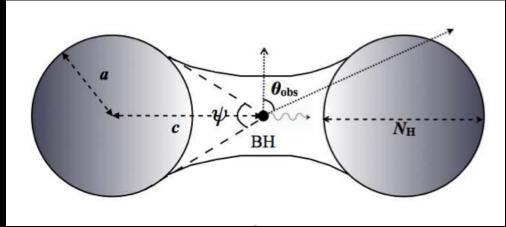


Image from http://www.mytorus.com

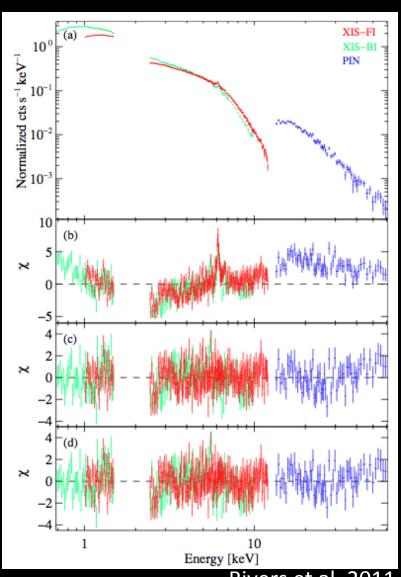
MYTorus model: torus of uniform density with half-opening angle of 60°. Self-consistently models:

- Fe line emission
- Compton reflection hump
- Absorption

MYTorus Results

- $\Gamma = 1.70 \pm 0.01$
- $A_{Fe} = 0.75 \pm 0.14$
- $N_{H,Tor} = 3.6_{-0.8}^{+1.3} \times 10^{24} \text{ cm}^{-2}$

Compton thick torus out of the line of sight



Rivers et al. 2011



Summary

- MCG-2-58-22 has no significant broad Fe line; narrow line is commensurate with the BLR
- Strong reflection component can be adequately modeled by the new MYTorus model
- Fe line EW is consistent with the amount of Compton-thick material
- For details see Rivers et al. 2011 ApJ 732, 36